

## Partiview Examples

in the `partiview/partiview/examples` folder

### hipmotion

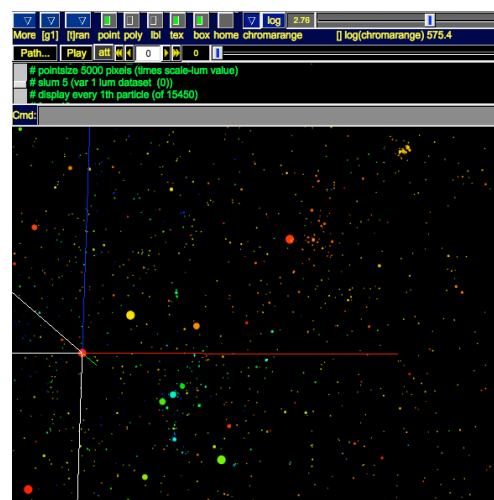


Nearby stars (from HIPPARCOS catalog) with space motions. Press the `>>` button to see the stars move at 10,000 years per second (or adjust the speed slider). Scripts and data are in the `data` folder – see “hipmotion”, which reads “hipbrightv.speck”.

In the `hipmotion` script, note the “`lum`”, “`color`”, “`cmap`” commands etc. to determine brightness/color; “`polylumvar`” and “`polysize`” to associate a screen-facing polygon with each star, proportional to its brightness; and “`texturevar`” and “`texture`” to apply a glow to each star. (Adjust brightness (“`slum`”) slider, or click “`tx`” texture button off.)

Also note the “`warp`” command. “`warp -extrap ...`” changes the position of each star as a (linear, in this case) function of time, based on its initial position and a group of three data variables giving velocity.

### hipchroma



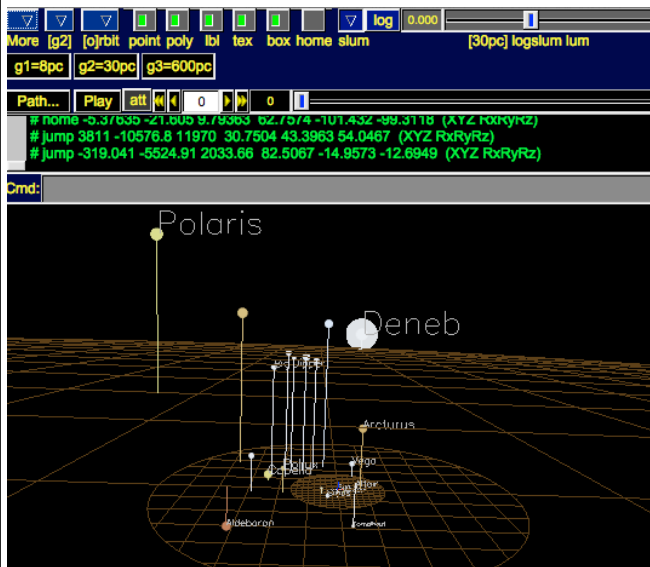
Nearby stars, color-coded by distance from the viewpoint. (Stars in a cluster look the same color!) Scripts/data are in the `data` folder – see “hipchroma”. Thanks to Prof. Anthony Fairall for the idea.

Designed for Chromatek stereo glasses, which make red things look nearby, blue ones distant. But it's helpful even without those glasses.

Try adjusting distance-to-color mapping: click the “`slum`” pulldown-menu at left of the brightness slider to make slider control other things. Choose “`chromarange`”, then drag the slider.

Try “`chromadepth off`” and “`chromadepth on`”.

## starwalk



If you could walk to the nearest star... Scripts/data in `starwalk` folder, starting with `starwalk.cf`. It defines three “groups”: “8pc”, “30pc” and “600pc”, also known as `g1`, `g2`, `g3`. Partiview creates a button for each – you can switch them separately on and off. It starts off with just the “8pc” group of nearest stars visible.

Note the label text (inside the `starwalk/stars8.cf`, `.speck`, etc. files), and the grid mesh and vertical spikes for each star, made from partiview “mesh” objects, generated by a program – the perl script “`speckv2galmap`”.

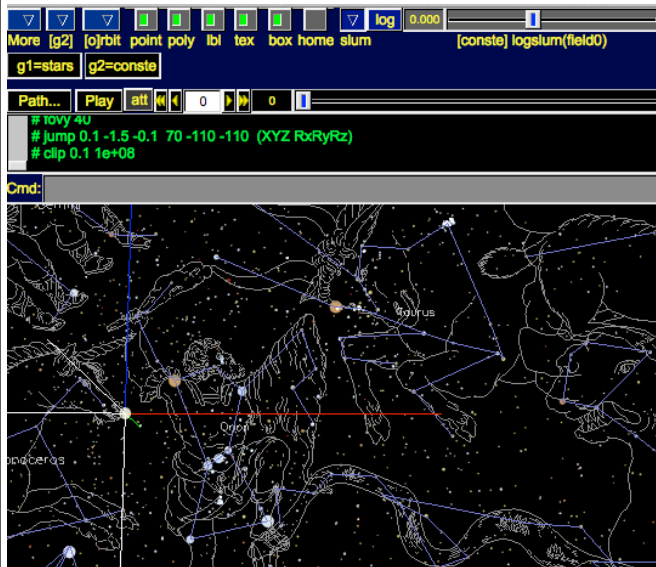
As seen from Earth, three stars – Deneb, Vega and Altair – look about equally bright. (Try flying to the origin crosshair, or click the “home” button.) But from elsewhere, we see that Deneb is a luminous supergiant – it's impressively bright over a large area, while Vega and Altair look small if we draw back a few tens of parsecs.

Star disk areas (`polylumvar ... area`) are proportional to each star's intrinsic luminosity, which makes their apparent brightness fade in the proper way with distance.

See also the writeup in: [starwalk/handout.pdf](#)

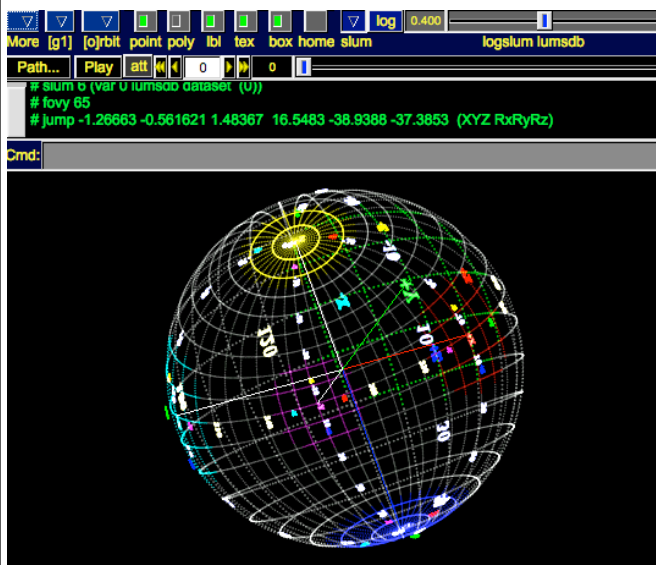
Guy Ottewell's book, *Astronomical Companion*, inspired this example.

## constellations



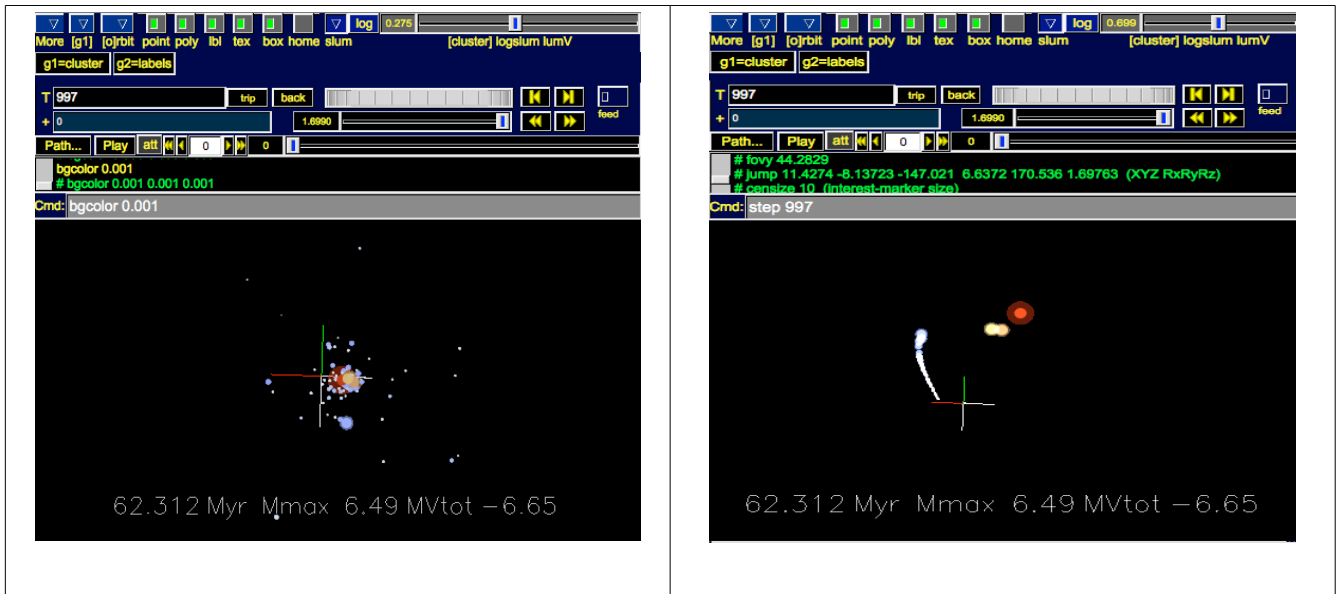
Pretty constellation pictures. Provided by Toshiyuki Takahei, National Observatory of Japan.

## spheregrid



Handy spherical marker, with text made entirely of dots. Demonstrates `.sdb` binary particle file format. See `spheregrid/mkspheregrid` for the perl script which created the data.

## cluster



An evolving cluster of stars. Two views of the same data: on the left, the stars' current positions; On the right, the “Hertzsprung-Russell” diagram, which plots the same stars on a 2-D graph of temperature (color) vs. luminosity. Press the >> key to watch the cluster evolve through 400 million years.

Uses the “warp” command to remap the positions of stars. Use “warp on” in the command box to see the temperature-luminosity view, or “warp off” to switch to the 3-D spatial position view. Used this way, “warp” lets you replace any particle's coordinates with any linear combinations of its data attributes.

Uses the .pb binary particle format, with each timestep in its own file, and a .speck file (cluster/clu8f/stars.speck) which loads each one into its corresponding partiview timestep (pb -t ...).

Also uses a second group to carry labels. The labels are made to stick to the screen, rather than moving in 3-D when you move the viewpoint, with the special “tfm camera” notation.

Stellar evolution calculated with Bill Paxton's wonderful “EZ” code, and cluster dynamics with Peter Teuben's “NEMO”.